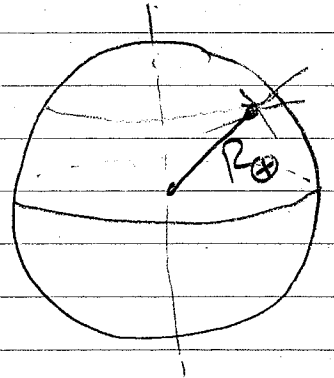


## Effects on the Earth

- Our reference frame on the Earth's surface

- As above, we choose an origin at  $\vec{R}$  given by  $R = R_0$ , polar + azimuthal angles  $\theta, \phi$



- +  $\phi$  can be chosen as longitude (measured from prime meridian)

- +  $\theta$  is colatitude, latitude  $\lambda$  is polar angle measured from equator  $\lambda = \pi/2 - \theta$

- + There are fixed axes  $\hat{i}, \hat{j}, \hat{k}$  w.r.t center of Earth.  $\hat{k}$  points north.

- Earth rotates with angular velocity  $\vec{\omega} = \omega \hat{k}$  where  $\omega = 2\pi / (1 \text{ day})$

- + The way we defined  $\theta, \phi$  as colatitude + longitude means  $\hat{r}, \hat{\theta}, \hat{\phi}$  axes are rotating

- + We can also set  $\hat{x}, \hat{y}, \hat{z}$  axes with origin at  $\vec{R}$  (rotating). Commonly,  $\hat{z}$  will be vertical,  $\hat{x} + \hat{y}$  along east + north.

• The force on an object as seen from a rotating frame is

$$\vec{F}_{\text{phys}} = 2m\vec{\omega} \times \dot{\vec{r}} - m\vec{\omega} \times (\vec{\omega} \times (\vec{R} + \vec{r}))$$

- + We are ignoring the transverse force since  $|\dot{\vec{\omega}}| \ll \omega^2$

- + For motion near the surface of the earth, i.e. where  $\vec{F}_{\text{grav}} \approx m\vec{g}$ , centrifugal force  $\approx -m\vec{\omega} \times (\vec{\omega} \times \vec{R})$

- Apparent gravity: which way is up?

- The force (real + fictitious) on a stationary object is

$$\vec{F} = -mg\hat{r} + m\omega^2 R_0 \hat{r} - m(\vec{\omega} \cdot \vec{r}) \vec{\omega}$$

+ The force of gravity is  $-GMm / (R_0 + r)^2 \hat{r} \approx -\left(\frac{GM}{R_0^2}\right) m \hat{r}$   
 so "gravitational" acceleration  $g \equiv GM/R_0^2$

+ Using  $\hat{k} = \cos\theta \hat{r} - \sin\theta \hat{\theta}$ , the centrifugal force is  
 $m\omega^2 R_0^2 [(1 - \cos^2\theta) \hat{r} + \cos\theta \sin\theta \hat{\theta}]$   
 $= m\omega^2 R_0^2 (\sin^2\theta \hat{r} + \frac{1}{2} \sin(2\theta) \hat{\theta})$

• A plumb line is hung to determine vertical (this is basically a pendulum at equilibrium).

+++ It hangs in the direction of the force,  $\vec{F}$ ,  
 an angle  $\alpha$  from  $\hat{r}$  where

$$\tan\alpha = \frac{\omega^2 R_0^2 \sin(2\theta)/2}{g - \omega^2 R_0^2 \sin^2\theta} \approx \frac{\omega^2 R_0^2 \sin(2\theta)}{2g} \approx \alpha$$

This is about 6 arc-min <sup>max</sup> (1 arc-min =  $\frac{1}{60}$  degree),  
 so small angle approx. is justified

+ Angle is toward the equator from  $\hat{r}$

+ We will take the plumb line as the direction of  $\hat{k}$ ,  
 our rotating Cartesian axis

• We can therefore define an apparent gravitational  
acceleration

$$\vec{g}_{\text{eff}} = \vec{g} + \vec{\omega} \times (\vec{\omega} \times \vec{r})$$

+ The direction is  $-\hat{k}$  (is downward vertical)

+ Magnitude is given by

$$g_{\text{eff}}^2 = (g - \omega^2 R_0^2 \sin^2\theta)^2 + \left(\frac{\omega^2 R_0^2}{2}\right)^2 \sin^2(2\theta)$$

or  $g_{\text{eff}} \approx g - \omega^2 R_0^2 \sin^2\theta$

+ At the pole,  $g_{\text{eff}} = g$ . At the equator,  $g_{\text{eff}} = g - \omega^2 R_0^2$   
 The difference is  $\approx 34 \text{ mm/s}^2$  or a few % of  $g$

+ Use  $\vec{g}_{\text{eff}}$  for  $\vec{g}$  in these notes. For convenience, we also shift colatitude  $\theta$  so  $\hat{k} \cdot \hat{z} = \cos \theta$ .

• In reality:

+ The earth's surface is roughly  $\perp \vec{g}_{\text{eff}}$   
So it is oblate, not spherical. Poles are closer to center, so  $g_{\text{eff}}$  is even slightly larger. (maybe more later)

+ Other deformations of Earth, like mountains, change  $\vec{g}_{\text{eff}}$  locally. + altitude

+ The standard value  $g \approx 9.806 \text{ m/s}^2$  is agreed upon by  $g_{\text{eff}}$  at one location theoretically adjusted to sea level + latitude  $45^\circ$ .

+ Ranges from  $\sim 9.78 \text{ m/s}^2$  to  $\sim 9.83 \text{ m/s}^2$ .

## - Coriolis Force Effects (by examples)

• Trade Winds: Warm air rises near equator + heads to poles. Cool air flows toward equator. The Coriolis force deflects it west, leading to trade winds. (At higher latitudes, other meteorological effects push winds the other direction - the case in most of US + Canada)

• Freely falling object

+ Take  $\hat{z}$  vertical,  $\hat{x}$  east,  $\hat{y}$  north

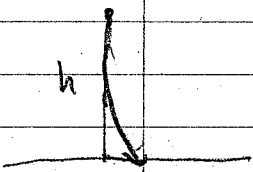
+ An object falls from  $x=0, y=0, z=h$ . Where does it land? The Coriolis force bends the path from vertical

+ We proceed perturbatively meaning we can treat Coriolis as a small effect. So we do not consider Coriolis forces due to velocity changes from Coriolis

+ With out the Coriolis force,  $x=y=0, z=h - \frac{1}{2}gt^2$   
Therefore, the Coriolis force is

$$-2m\vec{\omega} \times \dot{\vec{r}} \approx -2m\omega \hat{z} \sin \theta \dot{x} = +2m\omega g t \sin \theta \hat{x}$$

$$\left( \vec{\omega} = \omega (\sin \theta \hat{y} + \cos \theta \hat{z}) \right)$$



+ We therefore have  $\ddot{y} = 0$ , so there is no motion north, and

$$\ddot{x} = 2\omega g t \sin\theta \Rightarrow \dot{x} = \omega g t^2 \sin\theta$$

$$\Rightarrow x = \frac{1}{3} \omega g t^3 \sin\theta$$

+ Finally, we know it hits ground at  $t = \sqrt{2h/g}$ , so

$$x = \frac{1}{3} \omega \sqrt{8h^3/g} \sin\theta$$

At  $45^\circ$  latitude, this is  $\approx 16$  mm for  $h = 100$  m (small!)

+ Note: By definition, the apparent gravitational force is always toward  $-\hat{z}$ .

+ This is purely due to rotation, not the curvature of the earth, etc. It is independent of  $R_0$  as long as  $h \ll R_0$ .

◦ (Counter-) Example: Freely falling object in inertial frame.

+ Set up inertial unit vectors,  $\hat{x}'$ ,  $\hat{y}'$ ,  $\hat{z}'$  with the object initially at  $x' = y' = 0$ ,  $z' = h$  (and moving with the earth, i.e.  $\dot{x}' = \omega(R_0 + h) \sin\theta$ ).

+ We are setting  $\hat{z}'$  in the instantaneous  $\hat{r}$  direction for simplicity. Corrections for apparent vertical cancel (I think).

+ Subtlety: gravitational force is toward center of earth - we can't quite use uniform gravity. Using  $R_0 \gg x', z'$ , we have

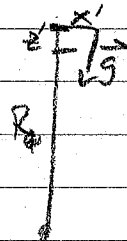
$$m\ddot{z}' = -mg, \quad m\ddot{x}' = -mgx'/R_0$$

+ We have  $z' = h - \frac{1}{2}gt^2$ ,  $x' = A \sin(\sqrt{g/R_0}t)$

$$\text{with } A = \omega(R_0 + h) \sin\theta \sqrt{R_0/g}$$

with  $t \leq \sqrt{2h/g} \ll \sqrt{R_0/g}$ ,  $x' \approx A \left[ \sqrt{g/R_0}t - \frac{1}{6}(\sqrt{g/R_0}t)^3 + \dots \right]$

must take into account b/c  $x'$  traveled is  $\propto R_0$



+ We now want extra  $x'$  distance object travels vs ground, which has  $x'$  speed  $\omega R \sin \theta$ . We have

$$\Delta x' \approx \omega h \sin \theta + \frac{1}{6} \omega R \sin \theta \left( \frac{g}{R \omega} \right)^3$$

When it hits the ground,  $t = \sqrt{2h/g}$

$$\Delta x' \approx \frac{1}{3} \omega \sin \theta \sqrt{8h^3/g}$$

+ We have made basically the same approximation (small deviation from falling along  $z/z'$  with uniform gravity) and gotten same answer. Have to be more careful to account for everything here.

+ Notes on reading: KT3 makes incorrect statement about the inertial frame version. They say the object starts with same  $v_{x'}$  as ground and speeds up as it falls due to angular momentum conservation. It actually starts faster b/c at greater height and slows down b/c  $z'$  velocity enters  $\vec{J}$ . Also, TMT does inertial frame calculation using an elliptical orbit. Correct but needlessly complicated.



• Foucault Pendulum: This is a pendulum that can move in both  $x$  and  $y$  directions freely. Oscillates quickly along a line that rotates slowly in  $xy$  plane.

+ We take the small angle approximation, so gravity + tension act as restoring forces

$$F_z = -mg + T \cos \alpha, \quad F_{xy} = -T \sin \alpha \begin{cases} \cos \phi \\ \sin \phi \end{cases}$$

At small angles,

$$\alpha \approx \sin \alpha \approx \frac{\sqrt{x^2 + y^2}}{l}, \quad \cos \alpha \approx 1, \quad T \approx mg$$

b/c the pendulum moves on the surface of a sphere.

+ We also have  $\vec{\omega} = \omega (\sin \theta \hat{y} + \cos \theta \hat{z})$ , so

the Coriolis force is

$$-2m \vec{\omega} \times \dot{\vec{r}} = +2m \omega \sin \theta \dot{x} \hat{z} - 2m \omega \cos \theta \dot{x} \hat{y} + 2m \omega \cos \theta \dot{y} \hat{x}$$

We can ignore Coriolis in  $\hat{z}$  direction b/c it just changes  $T$  a small amount. (args to  $\theta$  over fast oscillation)

+ In total, we find equations of motion

$$\ddot{x} = -gx/l + 2\omega \cos\theta \dot{y}, \quad \ddot{y} = -gy/l - 2\omega \cos\theta \dot{x}$$

To solve, set  $u = x + iy$ .

$$\ddot{u} + i2\omega \cos\theta \dot{u} + (g/l)u = 0$$

Define  $\omega_0^2 = g/l$ ,  $\omega_1 = \omega \cos\theta$ .

+ Guess a solution  $u = A e^{i\Omega t}$ . Then  $-\Omega^2 - 2i\omega_1\Omega + \omega_0^2 = 0$   
 $\Rightarrow \Omega = -i\omega_1 \pm \sqrt{\omega_1^2 + \omega_0^2} \approx -i\omega_1 \pm \omega_0$

That is

$$u = A e^{-i\omega_1 t} e^{i\omega_0 t} + B e^{-i\omega_1 t} e^{-i\omega_0 t}$$

T.T.  $= e^{-i\omega_1 t} [A e^{i\omega_0 t} + B e^{-i\omega_0 t}]$

+ The factor in square brackets is set by init. cond.

Suppose it is motion in  $x$ , so  $A = B = \text{real}$ .

The  $e^{-i\omega_1 t}$  factor then slowly rotates it toward the  $y$  axis.

+ Can understand from meridian frame if it is at the pole. The pendulum just oscillates in a line while the Earth rotates underneath.